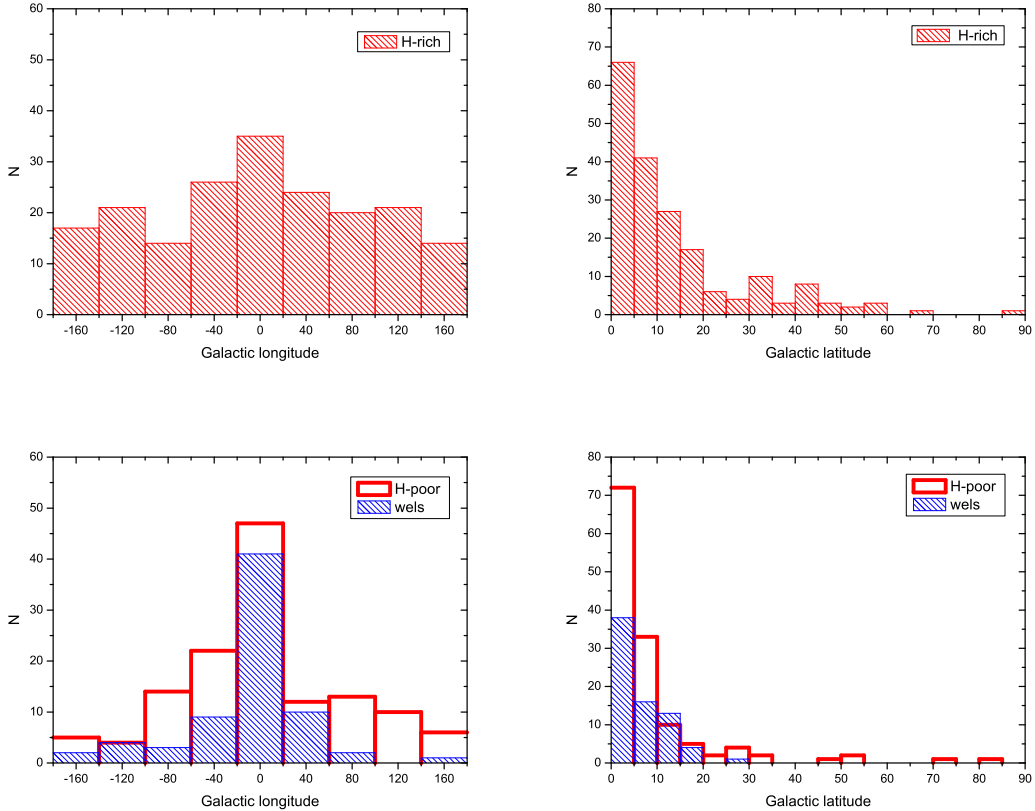

Fig. 1. Distributions in Galactic longitude and latitude of CSPN (of true and possible PN) that belong to H-rich, H-poor, and wels star groups. Note that H-poor PN are more concentrated towards the Galactic center than H-rich ones. The similarity between the wels and H-poor distributions is also noticeable.