

Statistical simulations of the dust foreground to cosmic microwave background polarization (Corrigendum)

F. Vansyngel¹, F. Boulanger¹, T. Ghosh^{1,2}, B. Wandelt^{3,4}, J. Aumont¹, A. Bracco^{1,5}, F. Levrier⁶,
P. G. Martin⁷, and L. Montier⁸

¹ Institut d'Astrophysique Spatiale, CNRS, Université Paris-Sud, Université Paris-Saclay, Bât. 121, 91405 Orsay Cedex, France

² California Institute of Technology, Pasadena, CA 91125, USA

³ Sorbonne Universités, UPMC Université Paris 6 et CNRS, UMR 7095, Institut d'Astrophysique de Paris, 98 bis bd Arago, 75014 Paris, France

⁴ Sorbonne Universités, Institut Lagrange de Paris (ILP), 98 bis Boulevard Arago, 75014 Paris, France

⁵ Laboratoire AIM, IRFU/Service d'Astrophysique, CEA/DSM, CNRS, Université Paris Diderot, Bât. 709, CEA-Saclay, 91191 Gif-sur-Yvette Cedex, France

⁶ LERMA, Observatoire de Paris, PSL Research University, CNRS, Sorbonne Universités, UPMC Université Paris 06, Ecole normale supérieure, 75005, Paris, France

⁷ CITA, University of Toronto, 60 St. George St., Toronto, ON M5S 3H8, Canada

⁸ CNRS, IRAP, 9 Av. Colonel Roche, BP 44346, 31028 Toulouse Cedex 4, France

A&A, 603, A62 (2017), <https://doi.org/10.1051/0004-6361/201629992>

Key words. polarization – ISM: general – comic background radiation – submillimeter: ISM – errata, addenda

The caption of Fig. 8 had a confusing description of the material shown in the plots. The correct caption together with the figure is given below.

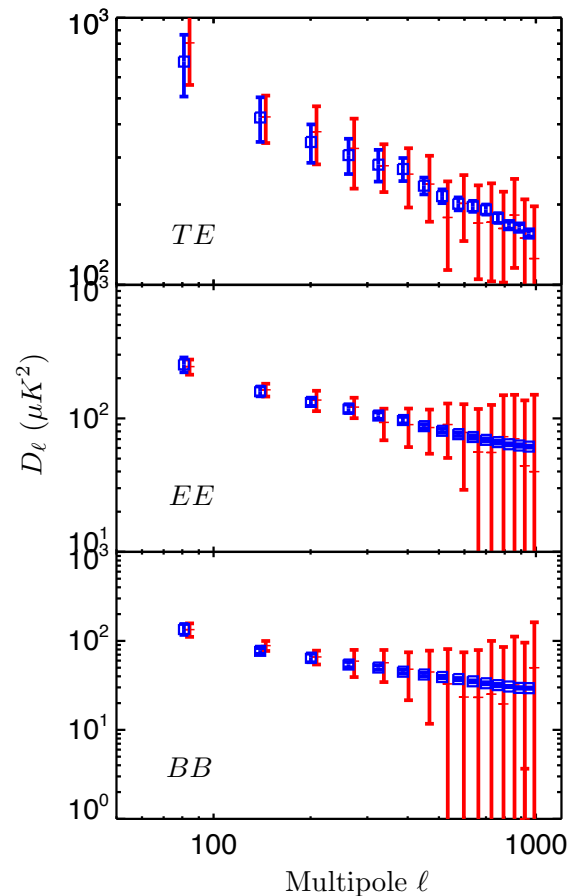


Fig. 8. High-resolution power spectra on the LR63 region; simulation vs. data. *From top to bottom:* *TE*, *EE*, and *BB* power spectra of the *Planck* 353 GHz, CMB-corrected maps (red crosses), and one high-resolution ($N_{\text{side}} = 2048$, $FWHM = 10'$) realization of the model (blue squares).